

lucheĭ v ideal'nykh kristallakh (Dynamical Scattering of X Rays in Ideal Crystals, Nauka, 1974 [Springer, 1978]).

²C. I. Iveronova and G. P. Revkevich, Teoriya rasseyaniya rentgenovskikh lucheĭ (Theory of X-Ray Scattering), Izd. MGU, 1978.

³J. Cowley, Diffraction Physics [Russ. transl.], Mir, 1979.

⁴V. A. Baskakov and B. Ya. Zel'dovich, Focusing of X-Rays via Multiwave Dynamic Diffraction, FIAN Preprint No. 191, 1978.

Translated by J. G. Adashko

Kinetic phenomena in the flow of a strongly rarefied molecular gas in an external field

V. D. Borman, S. Yu. Krylov, B. I. Nikolaev, V. A. Ryabov, and V. I. Troyan

Moscow Engineering Physics Institute

(Submitted 1 April 1980)

Zh. Eksp. Teor. Fiz. 79, 1787-1796 (November 1980)

A kinetic theory is developed for the effects that arise when a free-molecular polyatomic gas flows between two surfaces in an external field. The influence of the field on the transport processes is due to the nonequilibrium polarization of the gas molecules when they are nonspherically scattered from the surface of a solid, and to the destruction of this polarization in the field. The change of the gas flow velocity in a channel in a magnetic field, and the onset of a transverse heat flux between the surfaces (whose temperatures are equal) is examined in detail. In contrast to the previously investigated thermomagnetic phenomena, the considered effects in a gas stream can occur when the molecules are scattered from the surface not only inelastically but also elastically. At the same time, these effects occur only if the interaction with the surface is such that the states of the molecule before and after the collision are correlated.

PACS numbers: 47.45.Dt, 47.60.+i, 47.10.+g

1. INTRODUCTION

The influence of a magnetic field on heat flow in a strongly rarefied ($l \gg L$, where l is the molecule mean free path and L is the characteristic dimension) polyatomic gas (the thermomagnetic effect) has already been observed and investigated earlier.^{1,2} Other possible effects in an inhomogeneously heating gas in a magnetic field were also analyzed, such as the appearance of transverse heat and mass fluxes in a gas contained between two surfaces having different temperatures, or of thermomagnetic forces acting on the walls.³ The physical causes common to the changes in the transport processes in a magnetic field are the polarization of the molecules inelastically scattered from the solid surface and the precession of the magnetic moment of the molecule about the field direction. A distinguishing feature of the foregoing effects is the oscillatory character of the dependence of the macroscopic fluxes in the gas on the intensity of the constant external field at a fixed geometry of the problem. The concrete dependence of the macroscopic quantities on the intensity and orientation of the field is determined entirely by the law of nonspherical scattering of molecules by walls. Therefore the kinetic effects in a strongly rarefied gas in an external field serve as a unique source of information on the physical mechanism of the orientation-dependent interaction between molecules and the surface of a solid, and on the properties of the surface itself.^{2,4,5}

By virtue of the isotropy of the distribution of the molecules of the equilibrium gas with respect to their orientations and directions of motion, polarization of

molecules deflected from the surface can occur only in a nonequilibrium gas. The effects listed above are due to the temperature inhomogeneity of the system. It can be assumed that the molecules reflected (elastically and inelastically) from the walls become polarized also in the case of gas flow. The presence of a predominant direction of the velocity of the molecules incident on the surface and the dependence of the probability of the scattering on the mutual orientation of the velocity \mathbf{v} and of the angular momentum \mathbf{M} of the molecule should make the distribution function dependent also on the orientation of the vector \mathbf{M} , i.e., should lead to polarization of the molecules. The molecule precession produced when the external field is turned on changes this dependence (it destroys partially the polarization). As a result, the kinetic properties of the system are altered in an external field; in particular, the scalar transport coefficients acquire a tensor character.

In this paper we construct a theory of the phenomena connected with the influence of an external field on the transport processes in a stream of strongly rarefied polyatomic gas. We solve the problem of the flow of collisionless gas in a channel made up of two infinite surfaces in a magnetic field. We investigate the change of the channel resistance in the field and the onset of heat flow between the surfaces (which have equal temperatures). These effects are the Knudsen analogs of the known viscomagnetic effect⁶ and of the effect of viscomagnetic heat flow,⁷ which take place if $l \lesssim L$. They are produced, however, by another physical mechanism, namely polarization of the molecules by nonspherical scattering from the surface.

The considered phenomena have a polarization-precession mechanism and are not sensitive to the concrete cause of the precession of the molecules. Therefore the expression obtained below for the non-paramagnetic molecules in a magnetic field can be easily generalized to the case of paramagnetic molecules, as well as the case of dipole molecules in an electric field.

We must emphasize two substantial differences between the physical mechanism responsible for the onset of the considered effects in a gas stream, and the mechanism of the previously investigated thermomagnetic effects. First, the field should influence the transport processes in a gas stream not only in inelastic but also in elastic scattering of molecules by wall, when the total energy of the molecule remains unchanged by the collision. Consequently, an investigation of these effects makes it possible to study additionally also the elastic interaction of polyatomic molecules with a surface. This possibility is of fundamental importance, since elastic scattering should depend only on the structure of the surface, while inelastic scattering should depend also on the spectrum of the surface phonons of the solid.

The other distinguishing feature of the considered effects is, as will be shown below, that they appear only if the scattering of the molecules by the surface is such that the states of the molecule before and after the collision are correlated. On the other hand, if the scattering law takes into account only the sticking of the molecule to the surface and its evaporation from it,⁸ then the field has no effect on the transport processes in a stream of a collisionless gas. It follows from this, in particular, that at sufficiently low temperatures (under conditions of strong physical adsorption) the considered effects (in contrast to the thermomagnetic ones) can vanish.

2. KINETIC THEORY OF FLOW OF COLLISIONLESS GAS IN A MAGNETIC FIELD

We consider the influence of a magnetic field \mathbf{H} on the flow of strongly rarefied gas ($l \gg L$) in a channel between two infinite surfaces described by the equations $z=0$ and $z=L$. Let the gas move along the x axis under the influence of a specified density gradient ∇n and (or) a specified temperature gradient ∇T . The kinetic equation for the distribution function f of a gas with rotational degrees of freedom in a magnetic field⁹ takes in this case the form

$$v_x \rho f + v_z \frac{\partial f}{\partial z} + \gamma [\mathbf{M} \times \mathbf{H}] \frac{\partial f}{\partial \mathbf{M}} = 0, \quad (1)$$

$$f = (1 + \rho x) f; \quad f = f(\mathbf{v}, \mathbf{M}, z), \quad \rho = \frac{\nabla n}{n} + \frac{\nabla T}{T} \left(\frac{E}{T} - c_p \right),$$

where $E = mv^2/2 + M^2/2I$ is the energy of the linear molecule, γ is the gyromagnetic ratio, and c_p is the specified heat. Equation (1) can be solved jointly with the system of the boundary conditions on the wall, which relate the distribution functions of the incident (f_{inc}) and reflected (f_{ref}) molecules, and which are written in the form²

$$f_{\text{ref}} = \hat{W} f_{\text{inc}} = \int_{\mathbf{v}' \cdot \mathbf{k} < 0} \frac{|\mathbf{v}' \cdot \mathbf{k}|}{|\mathbf{v} \cdot \mathbf{k}|} W(\Gamma' \rightarrow \Gamma) f_{\text{inc}}(\Gamma') d\Gamma'. \quad (2)$$

Here $\Gamma = (\mathbf{v}, \mathbf{M})$, $d\Gamma = d\mathbf{v} d\mathbf{M}$, \mathbf{k} is the normal vector to the surface and W is the probability density of the scattering.

We integrate Eq. (1) along the characteristics in the range $0 \leq z \leq L$ and, using relations of the type (2) in the resultant boundary terms, we can change over to an integral kinetic equation, which we express in the following operator form

$$f(z) = \Theta(v_x) S^+ \hat{P}^+ \hat{W}^+ f(0) + \Theta(-v_x) S^- \hat{P}^- \hat{W}^- f(L), \quad (3)$$

$$S^\pm(z) = \exp\left(-\frac{z v_x \rho}{\xi}\right), \quad S^-(z) = S^+(L-z),$$

$$\xi = |\mathbf{v} \cdot \mathbf{k}| = |v_x|, \quad \Theta(a) = 1 \quad (a > 0), \quad \Theta(a) = 0 \quad (a \leq 0). \quad (4)$$

The operators \hat{W}^+ and \hat{W}^- are determined by the general relation (2) and describe the scattering of the molecules respectively by the lower ($z=0$) and upper ($z=L$) surfaces. The operators \hat{P}^\pm in (3) describe the precession of the molecules in the field as they move along the trajectory and are determined by the expressions

$$\hat{P}^+ = \hat{P}(z \leftarrow 0), \quad \hat{P}^- = \hat{P}(z \leftarrow L),$$

$$\hat{P}(z \leftarrow z_0) f(\Gamma, z_0) = \sum_{lmko} d_{m0}^{(l)}(\theta_H) d_{mk}^{(l)}(\theta_H) Y_{lk} \left(\frac{\mathbf{M}}{M} \right) \times \exp\left[i(\sigma - k) \varphi_H + im \frac{\gamma H(z - z_0)}{v_x} \right] \int Y_{l\sigma} \left(\frac{\mathbf{M}}{M} \right) f(\Gamma, z_0) d\Omega_M, \quad (5)$$

where θ_H and φ_H are the spherical field-orientation angles. The first and second terms in the right-hand side of (3) describes the distributions of the molecules moving up ($v_x > 0$) and down ($v_x < 0$), respectively. The functions S^\pm in (3) take into account the macroscopic motion of the gas. In the case $\nabla n = 0$ and $\nabla T = 0$, when $S^+ = S^- = 1$, Eq. (3) reduces to the previously obtained¹⁰ integral kinetic equation for a gas at rest in a magnetic field.

The error in (3), which manifests itself in the divergence of S^\pm (4) as $\xi = |v_x| \rightarrow 0$, is due to the fact that the intermolecular collisions in the gas are not taken into account [the collision integral in (1) is set equal to zero]. This result corresponds to the known divergence in the problem of planar Poiseuille flow as $(L/l) \rightarrow 0$, when the gas velocity becomes infinite at a finite value of the pressure gradient. To obtain a finite value of the flow velocity, it must be recognized¹¹ that the molecules with sufficiently small ξ manage to collide with other molecules before they reach the walls. We therefore put, neglecting intermolecular collisions,

$$\xi \geq \xi_{\text{min}} \approx \langle v \rangle L/l \quad (6)$$

($\langle v \rangle$ is the average thermal velocity of the molecules), and integrate in (2), which defines the operators \hat{W}^\pm , not over the half-space $\mathbf{v}' \cdot \mathbf{k} < 0$, but in the region $\mathbf{v}' \cdot \mathbf{k} < -\langle v \rangle L/l$. As will be indicated below, the assumption (6) has no effect whatever on the dependences of the macroscopic fluxes on the magnetic field. The validity of the last statement is quite obvious, however, from general considerations, inasmuch as in accordance with the precession mechanism a substantial contribution to the effect should be made only by molecules with $\xi \sim \langle v \rangle$.

We construct the solution of Eq. (3) by an iteration method, defining the expression for f in the n -th approximation as follows:

$$f^{(n)} = \Theta(v_z) S^+ \hat{P}^+ \hat{W}^+ f^{(n-1)}(0) + \Theta(-v_z) S^- \hat{P}^- \hat{W}^- f^{(n-1)}(L), \quad n=1, 2, 3, \dots \quad (7)$$

In accordance with (7), each iteration step corresponds physically to a refinement of the distribution function by taking into account one collision of the molecule with the surface. As the initial function $f^{(0)}$ we choose the Maxwellian distribution f_M . In accord with (1), the gas distribution function f is described here by the local Maxwellian distribution corresponding to specified macroscopic gradients.

We consider first the distribution f^+ of molecules moving upward ($v_z > 0$). After an arbitrary k -th iteration we have

$$f^{(k)+} = S^+ \hat{P}^+ \hat{W}^+ [S^- \hat{P}^- \hat{W}^-]_{z=0} [S^+ \hat{P}^+ \hat{W}^+]_{z=L} \dots [S^+ \hat{P}^+ \hat{W}^+]_{z=L} f_M. \quad (8)$$

The number of combinations $S^+ \hat{P}^+ \hat{W}^+$ in the right-hand side of (8) is equal to the number of the iteration k , and the first position on the right is occupied by $S^+ \hat{P}^+ \hat{W}^+$ taken at $z=L$, if k is odd, or by $S^- \hat{P}^- \hat{W}^-$ with $z=0$ if k is even. Next, the operators \hat{P} and \hat{W} in the first combination $S^+ \hat{P}^+ \hat{W}^+$ on the right should be left out, since $\hat{W} f_M = f_M$ (the known conservation of the Maxwellian distribution), and \hat{P} is a unit operator in the subspace of functions that do not depend on the orientation of \mathbf{M} .

For the analysis that follows, we represent the scattering probability in the form of the sum

$$W = W_0 + \varepsilon W',$$

$$W' = W_1(v', M' \rightarrow v, M) + W_2(v', M' \rightarrow v, M^2) + W_3(v', M' \rightarrow v, M), \quad (9)$$

where W_0 does not depend on the orientation of the angular momentum of the molecule, and the functions W_1 , W_2 , and W_3 describe the nonspherical interaction with the surface. The nonsphericity parameter ε will be assumed small (the experimental value of ε in inelastic scattering, obtained from an investigation of the thermomagnetic effect,⁴ is given by $\varepsilon^2 \sim 10^{-3} - 10^{-2}$).

To determine the field dependence of the macroscopic currents in the gas (which are not connected with the transport of \mathbf{M}) it is necessary to calculate that part f_H of the distribution function which depends on \mathbf{H} and is isotropic in \mathbf{M} . When (9) is taken into account, expression (8) leads to a nonzero result for $f_H^{(k)}$ only in the second and higher-order approximations in ε and at $k \geq 3$

$$f_H^{(k)+} = \varepsilon^2 \sum_{z=0}^{k-2} S^+ \hat{W}_0^+ [S^- \hat{W}_0^-]_{z=0} \dots [S^+ \hat{W}_0^+]_{z=L} [S^+ \hat{P}^+ \hat{W}_1^+]_{z=0} [S^+ \hat{P}^+ \hat{W}_1^+]_{z=L} + O(\varepsilon^3). \quad (10)$$

The summation is carried out here over all the permutations of the combination $\hat{W}_2 S^+ \hat{P}^+ \hat{W}_1$ in the written-out product containing $k-3$ operators \hat{W}_0 and one operator \hat{W}_1 and \hat{W}_2 each [the explicit forms of the operators \hat{W}_0 , W_1 , and W_2 follow from (2) when W is replaced respectively by W_0 , W_1 , and W_2]. Depending on the location of this combination in the product, the operator indices are either plus or minus.

The expression (10) reflects the physical mechanism

of the effect of the field on the transport processes, an explanation of which calls for allowance for at least two nonspherical collisions of the molecules with the walls. In the first of them the molecules become polarized (the polarization is specified in (10) by the operator \hat{W}_1), while precession of the molecules in the field partially destroys this polarization (the action of \hat{P}), and the second nonspherical collision (the operator \hat{W}_2) transforms the field-dependent distribution, which is anisotropic in \mathbf{M} , into an isotropic distribution. The dependence of the latter on \mathbf{H} describes the change of the considered transport processes in the field.

We assume hereafter that the density and temperature gradients are small ($\nabla n \ll n/l$, $\nabla T \ll T/l$) and confine ourselves to an approximation linear in ∇n and ∇T . We must then leave out of (10) all but one of the functions S . Taking into account the relations $\hat{W}_0 f_M = f_M$ and $\hat{W}_1 f_M = 0$, we obtain

$$f_H^{(k)+} = \varepsilon^2 \sum_{z=0}^{k-2} \sum_{z=0}^{k-1} \hat{W}_0^+ \hat{W}_0^- \dots \hat{W}_z^+ \hat{P}^+ \hat{W}_1^+ \hat{W}_0^+ \dots \hat{W}_0^+ \frac{v_z \rho L}{\zeta} f_M. \quad (11)$$

The second summation in (11) implies sifting through all the possible operators W_0 on the right of the combination $\hat{W}_2 P \hat{W}_1$. This summation is the result of the fact that it is possible to retain in (10) any function S from among those on the right of the indicated operator combination.

It is easy to obtain in similar fashion an expression for the distribution $f_H^{(k)-}$ of the molecules moving down ($v_z < 0$); this expression differs from (11) in only the interchange of the plus and minus indices.

We assume next that the surfaces are macroscopically isotropic (e.g., polycrystalline), i.e., there is no physically preferred direction in the plane of the surface. Then, as can be easily verified, the functions B_1 and B_2 calculated in accordance with the equations

$$\hat{W}_0 \hat{W}_0 \dots \hat{W}_0 A_1(v^2, M^2, \zeta) v_i = B_1(v^2, M^2, \zeta) v_i; \quad i=x, y, z; \quad (12)$$

$$\int d\Gamma A_2(v^2, M^2, \zeta) v_i \hat{W}_0 \hat{W}_0 \dots \hat{W}_0 D(v, M^2) = \int d\Gamma B_2(v^2, M^2, \zeta) v_i D(v, M^2),$$

turn out to be independent of the tangential component of the velocity if the functions A_1 and A_2 are independent of it, and the function D is arbitrary.

Letting the number of iterations k in (11) go to infinity, and using relation (12) as well as the explicit forms of the operators \hat{P} , \hat{W}_1 , and \hat{W}_2 , it is easy to obtain an expression for the field-dependent part of the macroscopic flux of the quantity A ($A=1$, E , or mv_x if the flux referred to is that of the number of particles, of the heat, or of the normal momentum):

$$\begin{aligned} \langle Av_i \rangle_H &= \frac{1}{n} \int_{v_z > 0} Av_i f_H^+ d\Gamma + \frac{1}{n} \int_{v_z < 0} Av_i f_H^- d\Gamma \\ &= -\varepsilon^2 L \iiint_{v_x, v_y, v_z > \zeta_{min}} d\Gamma d\Gamma' d\Gamma'' \frac{v_i v_z'' f_M''}{\zeta} \left[\frac{\nabla n}{n} + \frac{\nabla T}{T} \left(\frac{E''}{T} - c_p \right) \right] B_1(v^2, M^2, \zeta) \\ &\quad \times B_2(v''^2, M''^2, \zeta'') \sum_{l=mv_0} (-1)^{l+m} d_{-mv}^{(l)}(\theta_H) d_{no}^{(l)}(\theta_H) \exp \left[i(v+\sigma)\varphi_H \right] \\ &\quad + im \frac{\gamma HL}{\zeta'} \left[W_{100}^+(\Gamma_R \Gamma) W_{000}^-(\Gamma'' \Gamma_R') \pm W_{100}^-(\Gamma' \Gamma_R) W_{000}^+(\Gamma_R'' \Gamma') \right]. \end{aligned} \quad (13)$$

Here

$$\Gamma = (\mathbf{v}, M^2), \quad d\Gamma = 4\pi d\mathbf{v} M dM, \quad \Gamma_n = (\mathbf{v}_n, M^2), \quad \mathbf{v}_n = \mathbf{v} - 2\mathbf{k}(\mathbf{v}\mathbf{k}), \quad (14)$$

$$W_{lmjn} = \iint Y_{lm} \left(\frac{\mathbf{M}'}{M'} \right) Y_{jn} \left(\frac{\mathbf{M}}{M} \right) W'(\mathbf{v}' \mathbf{M}' \rightarrow \mathbf{v}, \mathbf{M}) d\omega_{\mathbf{M}'} d\omega_{\mathbf{M}}.$$

In the last square bracket of (13) it is necessary to take the plus sign when calculating $\langle v_i \rangle$, $\langle E v_i \rangle$ ($i = x, y$) and $\langle v_x v_x \rangle$ and the minus sign for the calculation of $\langle E v_x \rangle$ and $\langle v_x v_i \rangle$ ($i = x, y$). The explicit forms of the functions B_1 and B_2 in (13) depends on A and on the probability W_0 of the spherically symmetrical scattering. All that matters to us hereafter is that B_1 and B_2 do not depend on the tangential component of the molecule velocity.

Expression (13) makes it possible to calculate the changes in the heat and mass fluxes of the gas in the field, and also the normal and tangential forces acting on the surfaces. As follows from (13), the character of the dependence of the macroscopic fluxes on \mathbf{H} is determined entirely by the nonspherical part of the scattering probability W' and does not depend on the explicit form of W_0 . The latter influences only the magnitude of the effects.

We shall use below, by way of example, a model expression for W' , describing inelastic scattering of the molecules N_2 and CO by surfaces of (polycrystalline) gold or platinum. This expression, deduced from an investigation of the thermomagnetic effect,² is obtained by retaining in the expansion

$$W'(\mathbf{v}', \mathbf{M}' \rightarrow \mathbf{v}, \mathbf{M}; \mathbf{k}) = |\mathbf{v}\mathbf{k}| e^{-\beta E'} \sum_s \beta_s A_s; \quad s = \{(l_1' l_1) j_1 (l_2' l_2) j_2 l_3\};$$

$$A_s = [Y_{l_1'}(\mathbf{v}') \times Y_{l_1}(\mathbf{v})]^{j_1} [Y_{l_2'}(\mathbf{M}') \times Y_{l_2}(\mathbf{M})]^{j_2} Y_{l_3}(\mathbf{k}) \quad (15)$$

only the first four terms:

$$\{(00)0(02)22\}, \{(00)0(20)22\}, \{(11)1(02)22\}, \{(11)1(20)22\}. \quad (16)$$

3. CHANGE OF THE CHANNEL RESISTANCE IN THE FIELD

Using the model of nonspherical scattering (15)–(16), calculation of the change of the gas flow velocity in the field by formula (14) leads to the following results at various orientations of \mathbf{H} :

$$\Delta \langle v_x \rangle = \langle v_x \rangle_{\mathbf{H}} - \langle v_x \rangle_{\mathbf{H}=0},$$

$$\Delta \langle v_x \rangle |_{\mathbf{H}=\mathbf{H}_n} = \Delta \langle v_x \rangle |_{\mathbf{H}=\mathbf{H}_n} = \psi_1 J_1(\omega\tau),$$

$$\Delta \langle v_x \rangle |_{\mathbf{H}=\mathbf{H}_v} = \psi_2 J_2(\omega\tau); \quad \psi_1 = \varepsilon^2 L \left(\frac{2T}{m} \right)^{1/2} \left(C_1 \frac{\sqrt{n}}{n} + C_2 \frac{\sqrt{T}}{T} \right). \quad (17)$$

Here

$$J_k(\omega\tau) = \int_0^\infty x^2 e^{-x^2} \left[1 - \cos \left(k \frac{\omega\tau}{x} \right) \right] dx, \quad k=1, 2, \quad (18)$$

$$\omega = \gamma H, \quad \tau = L(2T/m)^{-1/2}, \quad x = \xi'(2T/m)^{-1/2}.$$

The constants C_1 and C_2 (~ 1) can be calculated if we know the expansion parameters β_s in (15) which correspond to the terms (16).

The integration variable x in the integral (18) is a dimensionless normal component of the molecule velocity, therefore the lower limit of integration, in accord with (6), would have to be set equal to $\xi_{\min}(2T/m)^{-1/2}$. However, by virtue of the rapid oscillations of the integrand at small x , the lower limit can be set equal to zero with practically no effect on the value of J_k . The foregoing confirms the validity of the statement made above that assumption (6) does not influence the

field dependence of the macroscopic current.

As follows from (17), the gas velocity in the channel is determined by the parameter $\omega\tau$, which is equal to the product of the molecule precession frequency in the field by the characteristic time of their travel between the walls. The dependence of $\Delta \langle v_x \rangle$ on $\omega\tau$ is similar to the corresponding dependences of the thermomagnetic effect¹ and takes the form of damped oscillations. The first maximum is reached here at $\omega\tau \sim 1$, and $\Delta \langle v_x \rangle$ saturates as $\omega\tau \rightarrow \infty$. The magnitude of the effect at saturation is the same for the three different field orientations.

In accord with (17), the rate of flow of the gas in the channel decreases when the magnetic field is turned on. Since the gradients ∇n and ∇T are fixed, this means that the channel resistance decreases in the field. As shown by the analysis of (13), this result is typical of all those terms of the expansion (15) with $l_1' + l_1 + l_2' + l_2 + l_3$ an even number, i.e., when the probability of scattering by the surface is invariant to reversal of the vectors \mathbf{v}' , \mathbf{v} , \mathbf{M}' , \mathbf{M} , and \mathbf{k} .

The expansion terms with $l_1' + l_1 + l_2' + l_2 + l_3$ odd, on the contrary, decrease the resistance of the channel in the field. We note that this general result can be obtained simply from physical considerations by using arguments previously advanced^{2,3} to explain the sign of the thermomagnetic effect.

An important feature of the considered effect, compared with the thermomagnetic effects, is that the change of the channel resistance in a field can occur not only for inelastic but also for elastic reflection of the molecules from the surfaces. In the latter case the total energy of the molecule is not changed by the collision, and the coefficients β_s of the corresponding terms of the expansion (15) contain $\delta(E' - E)$ as a factor. The contributions of the elastic and inelastic collisions to the effect can be separated by investigating, for example, its temperature dependence. The possibility of obtaining information on elastic nonspherical scattering does not depend on the phonon spectrum of the solid and is determined only by the structure of the surface.

In accord with formula (17), in the case of inelastic scattering the relative change of the flow velocity in the field is of the order of the square of the nonsphericity parameter and amounts to $\Delta \langle v_x \rangle / \langle v_x \rangle \sim \varepsilon^2 \ln^{-1}(l/L) \approx 1\%$. We note that the nonsphericity of the elastic scattering may turn out to be stronger than for the elastic one. (Elastic scattering can apparently be treated as reflection from an unperturbed "smooth" surface, while inelastic scattering can be regarded as scattering by thermal surface fluctuations that can lead, by virtue of their stochastic character, to a weakening of the dependence on the molecule orientation.) The presented estimate of the magnitude of the effect is only the lower bound of the summary effect, which can reach large values in the presence of both elastic and inelastic scattering.

As seen from (13)–(15), contributions to the effect are made only by those terms of the expansion (15) in which the indices l_1 and l_2 or l_1' and l_2' are not simultaneously

equal to zero. In other words, a change in the resistance of the channel in the field takes place if the non-spherical scattering from both walls is such that the states of the molecules before and after the collision are correlated.

4. TRANSVERSE HEAT FLUX

Calculation of the heat flux $E\langle v_x \rangle$ with the aid of relations (13)–(16) yields

$$\langle E v_x \rangle|_{\mathbf{H} \parallel \mathbf{K}} = 0, \quad \langle E v_x \rangle|_{\mathbf{H} \perp \mathbf{K}} = \psi_2 \sin \varphi_H I_2(\omega\tau);$$

$$\psi_2 = \varepsilon^2 L T \left(\frac{2T}{m} \right)^{1/2} \left(C_3 \frac{\nabla n}{n} + C_4 \frac{\nabla T}{T} \right); \quad |C_3|, |C_4| \sim 1; \quad (19)$$

$$I_2(\omega\tau) = \int_0^{\omega\tau} x^2 e^{-x^2} \sin \left(2 \frac{\omega\tau}{x} \right) dx.$$

According to (19), energy transport is produced by turning on a field that is not collinear to the flow of gas between surfaces of equal temperature. The transverse heat flow is odd in the field and is an oscillating function of $\omega\tau$, reaching the first maximum at $\omega\tau \sim 1$ and tending to zero as $\omega\tau \rightarrow \infty$. The maximum heat flux in the flow of gas under the influence of, say, a temperature gradient is of the order of $\varepsilon^2 L (2T/m)^{1/2} \nabla T$. Just as in the preceding case, this estimate is a lower bound, inasmuch as allowance for the elastic scattering by the surfaces can increase the effect. In the case of pure elastic scattering of the molecules from both surfaces, however, this effect will not occur. To produce a transverse heat flux it is necessary that nonspherical scattering of the molecules from at least one of the walls be accompanied by energy exchange with the surface.

In contrast to the influence of the field on the gas flow, a transverse heat flux appears also if the scattering by one of the surface takes place without a correlation between the states of the molecule before and after the collision, i.e., it can be described with account taken of only the sticking and evaporation processes. There will be no effect, however, if this "correlationless" scattering is a characteristic of both surfaces (for example, under conditions of strong physical sorption).

We note that expression (19) coincides (accurate to within a factor of suitable dimensionality) with the expression for transverse thermomagnetic gas flow,³ as it should for crossover effects in accord with the Onsager principle.

5. OTHER KINETIC EFFECTS IN A MAGNETIC FIELD

We have considered above two kinetic phenomena that are of greatest interest from the point of view of the possibility of their experimental investigation. The complete list of the possible effect in a stream of a strongly rarefied gas, which are different physical manifestations of nonequilibrium polarization of the molecules and its destruction in the field, is much larger. A magnetic field not collinear with the gas flow lifts the degeneracy of the system with respect to directions perpendicular to the flow. As a result, turning-on a field can give rise also to heat and mass fluxes in a direction perpendicular to the main gas flow (along the y axis), as well as to normal and tangential forces that act on the walls. The dependence

TABLE I.

	Field orientation	Parity with respect to H	Parity with respect to K
$\langle E v_x \rangle$	$\begin{cases} 2 \\ 1 (*) \end{cases}$	-	+
$\Delta \langle v_x \rangle$	$\begin{cases} 1, 2, 3 \\ 3 (*) \end{cases}$	+	-
$\langle E v_y \rangle$	$\begin{cases} 1, 2, 3 (*) \\ 3 \end{cases}$	-	+
$\langle E v_z \rangle$	$\begin{cases} 1, 2, 3 (*) \\ 3 \end{cases}$	+	-
$\langle E v_x \rangle$	$\begin{cases} 1, 2, 3 (*) \\ 3 \end{cases}$	-	+

of all the macrofluxes in the gas on the magnitude and orientation of the field follows from the general relation (13).

Without presenting here the detailed results of the calculations, we note only that the possible existence of various macrofluxes in a field in some particular direction and their general properties (e.g., whether they are even or odd in the field) can be explained by starting from simple considerations of the spatial symmetry of the system.¹² To this end it is necessary to construct a phenomenological expression (of suitable tensor character) for the given macroscopic quantity out of the three characteristic vectors of the problem: the vector $\mathbf{u} = \langle \mathbf{v} \rangle / \langle v \rangle$ directed along the gas flow, the magnetic-field pseudovector \mathbf{H} , and the vector \mathbf{K} directed normal from the surface of one material to the surface of another material. Such analysis shows, in particular, that if the interaction of the molecules with the surfaces is not invariant to space reflection, then the vector fluxes in the system can be odd in \mathbf{K} , i.e., they can reverse sign when surfaces made of different materials exchange place. The results for vector macrofluxes that are odd in \mathbf{u} in a field are listed in the table, where the asterisk marks effects that arise when the scattering by the surface is invariant to inversion, and the numbers designate the different field orientations at which the effect takes place, namely: 1) $\mathbf{H} \parallel \mathbf{u}$, 2) $\mathbf{H} \parallel \mathbf{K} \times \mathbf{u}$, and 3) $\mathbf{H} \parallel \mathbf{K}$.

The kinetic theory constructed above can predict also the field dependence of the thermomolecular pressure difference (in the free-molecular limit). This effect consists of the onset of a pressure difference between two gas volumes connected by a sufficiently thin capillary and kept at different temperatures. It is easy to show that the dependence of the pressure difference on the field should coincide in form with the dependence of the gas-flow velocity described by expressions (13) and (17).

We note in conclusion that since the motion of the gas-mixture components in the free molecular regime is independent, the theory developed above can yield also without difficulty expressions for the diffusion and thermal-diffusion coefficients in a field.

The authors thank L. A. Maksimov for a discussion of the formulation of the problem and of the results.

¹V. D. Borman, L. A. Maksimov, B. I. Nikolaev and V. I. Troyan, Dokl. Akad. Nauk SSSR **207**, 1082 (1972) [Sov. Phys. Dokl. **17**, 990 (1973)]; Zh. Eksp. Teor. Fiz. **64**, 526 (1973) [Sov. Phys. JETP **37**, 269 (1973)].

²V. D. Borman, B. I. Buttsev, S. Ĭu. Krylov, B. I. Nikolaev and V. I. Troyan, Zh. Eksp. Teor. Fiz. **70**, 929 (1976) [Sov. Phys. JETP **43**, 484 (1976)].

- ³V. D. Borman, S. Yu. Krylov, L. A. Maksimov, B. I. Nikolayev and V. I. Troyan, *Phys. Lett.* **67A**, 25 (1978).
⁴V. D. Borman, S. Yu. Krylov, B. I. Nikolaev, V. A. Ryabov and V. I. Troyan, *Zh. Eksp. Teor. Fiz.* **71**, 1373 (1976) [*Sov. Phys. JETP* **44**, 719 (1976)].
⁵V. D. Borman, B. I. Buttsev, B. I. Kikolaev, A. P. Popov and V. I. Troyan, *Zh. Eksp. Teor. Fiz.* **77**, 2297 (1979) [*Sov. Phys. JETP* **50**, 1105 (1979)].
⁶J. J. M. Beenakker and F. R. McCourt, *Ann. Rev. Phys. Chem.* **21**, 47 (1970).
⁷G. E. J. Eggermont, P. W. Hermans, L. J. F. Hermans and J. J. M. Beenakker, *Phys. Lett.* **57A**, 29 (1976).
⁸S. Yu. Krylov, V. D. Borman, B. I. Nikolaev and V. I.

- Troyan, *Zh. Eksp. Teor. Fiz.* **67**, 2122 (1974) [*Sov. Phys. JETP* **40**, 1053 (1974)].
⁹Yu. Kagan and L. A. Maksimov, *Zh. Eksp. Teor. Fiz.* **41**, 842 (1961) [*Sov. Phys. JETP* **14**, 604 (1961)].
¹⁰V. D. Borman, S. Yu. Krylov and B. I. Nikolaev, *Zh. Eksp. Teor. Fiz.* **76**, 1551 (1979) [*Sov. Phys. JETP* **49**, 787 (1979)].
¹¹C. Cercignani, *Theory and Applications of Boltzmann's Equations, Teoriya i prelozheniya uravnenii Bolts' mana*, (Russ. Transl.) Mir, 1978, p. 341.
¹²V. D. Borman, S. Yu. Krylov, L. A. Maksimov, B. I. Nikolaev and V. I. Troyan, *Izv. AN SSSR*, 124 (1979).

Translated by J. G. Adashko

Closed equation for turbulent heat and mass transport

M. A. Vorotyntsev, S. A. Martem'yanov, and B. M. Grafov

Institute of Electronics, USSR Academy of Sciences

(Submitted 16 June 1980)

Zh. Eksp. Teor. Fiz. **79**, 1797-1808 (November 1980)

A tensor nonlocal relation is derived between the mass (heat) flux and the gradient of the average density (temperature) for a turbulently flowing liquid. The laws governing the turbulent mass transport in a diffuse boundary layer near a flat solid boundary are investigated, including the section where the layer thickness is constant. It is shown that at the start of the inlet section there is a region where the connection between the turbulent diffuse flow and gradient of the average density is essentially nonlocal. In the remaining part of the inlet section and in the region of the stabilized diffuse layer it is possible to obtain approximately a local relation between these quantities. The contribution made to the average diffusion flux on the surface by the hydrodynamic turbulent pulsations first decreases with increasing longitudinal coordinate, and then begins to increase. Longitudinal turbulent transport predominates up to the minimum point, and normal transport beyond this point. In the stabilized region, the turbulent diffusion coefficient takes on different functional forms at different distances from the surface.

PACS numbers: 47.25. - c, 47.25.Fj, 47.25.Jn, 47.10. + g

1. INTRODUCTION

In view of the wide prevalence of turbulent flows, the questions of heat and mass transport in turbulent streams attract much attention. From the theoretical point of view, the principal problem is the closing of the averaged transport equations: the density J_{turb} of the turbulent flux of matter or of heat must be connected with the distribution of the average density of the matter or the average temperature \bar{T} . As a rule, a local relation is assumed to exist between J_{turb} and the gradient of the average density or temperature¹⁻⁶:

$$J_{\text{turb}} = -D_{\text{turb}} \nabla \bar{c} \quad (1)$$

(to be specific, we discuss below the mass-transport problem).

The phenomenologically introduced turbulent-diffusion coefficient D_{turb} depends on the spatial coordinates, particularly on the distance to the solid surfaces. In some papers,⁹ several phenomenological quantities are introduced in the form of a tensor \bar{D}_{turb} that generalizes relation (1). To find the coefficient (or tensor) D_{turb} , it is customary to use the Reynolds analogy between D_{turb} and the turbulent viscosity coefficient ν_{turb} ¹⁻⁹:

$$D_{\text{turb}}(\mathbf{r}) \sim \nu_{\text{turb}}(\mathbf{r}). \quad (2)$$

For the last quantity, a power-law variation is usually

postulated near the boundaries of solids:

$$\nu_{\text{turb}} \propto y^k, \quad (3)$$

where a value 3 or 4 is assumed for the exponent k .

2. AVERAGED MASS-TRANSPORT EQUATIONS

The purpose of the present paper is to derive a relation between J_{turb} and $\nabla \bar{c}$ on the basis of the initial (non-averaged) equation of convective diffusion in incompressible liquids¹⁻⁷:

$$\partial c / \partial t + \mathbf{v} \nabla c = D \Delta c. \quad (4)$$

Here $\mathbf{v}(\mathbf{r}, t)$ is the instantaneous distribution of the velocities of the liquid, D is the molecular-diffusion coefficient, and $c(\mathbf{r}, t)$ is the field of the impurity densities. The latter is assumed to be too small to influence the hydrodynamic characteristics of the flow.

We resolve the velocity of the liquid and the density into averaged and pulsating components

$$\begin{aligned} \mathbf{v}(\mathbf{r}, t) &= \mathbf{u}(\mathbf{r}, t) + \mathbf{v}'(\mathbf{r}, t), & c(\mathbf{r}, t) &= \bar{c}(\mathbf{r}, t) + c'(\mathbf{r}, t), \\ \mathbf{u}(\mathbf{r}, t) &= \langle \mathbf{v}(\mathbf{r}, t) \rangle, & \bar{c}(\mathbf{r}, t) &= \langle c(\mathbf{r}, t) \rangle. \end{aligned} \quad (5)$$

Here and elsewhere, the brackets $\langle \dots \rangle$ denote averaging, while the prime denotes pulsating quantities.

Averaging of (4) leads to the fundamental equation of convective diffusion in turbulent flow: