

MEAN ENERGY OF THE Y^{90} BETA SPECTRUM

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The mean energy of the Y^{90} beta-ray spectrum was measured calorimetrically and the value 933 ± 18 kev obtained.

VARIOUS methods may be employed for an experimental determination of the average energy of the beta spectrum, \bar{E}_β . The calorimetric method, however, is at the present time the simplest and the most accurate. We measured calorimetrically the mean energy of the beta spectrum of Y^{90} , which is one of the isotopes most frequently used in radiometry and biology. The procedure was similar to that employed previously.^[1]

The Y^{90} source was obtained by irradiating in a reactor 1–1.5 g of yttrium oxide. The initial activity was usually on the order of 300–500 mC. The heating effect was measured using a double static calorimeter with a sensitivity of $\sim 2.5 \times 10^{-5}$ watt/mm. Measurements were made with four different Y^{90} sources. For each of these, calorimetric measurements lasted from three to seven days. The individual experimental points fell properly on straight lines which correspond to a decay with a 64.9 hour period. This is consistent with recent data on the half-life of Y^{90} (64.8 ± 0.2 hours)^[2] and is evidence of the purity of the sources; this was also controlled by measurements on a scintillation gamma spectrometer. Corrections were introduced into the calorimetric measurement results for the energy of internal bremsstrahlung of the Y^{90} absorbed in the calorimeter, and for the thermal inertia of the latter.^[3]

After the calorimetric measurements, all of the active yttrium was dissolved in weak nitric acid and the absolute activity of a known aliquot of the solution was measured by a 4π counter.

The mean energy of the Y^{90} beta-ray spectrum obtained when the results for all four sources were averaged was 933 ± 18 kev. The noticeable discrepancy between this value and the quantity $\bar{E}_\beta(Y^{90}) = 895 \pm 35$ kev obtained by Caswell^[4] using an extrapolation chamber may, it seems, be explained by the reduced value of ionization energy

ϵ ($\epsilon = 32.5$ ev) adopted by that author. If we recompute Caswell's results using the value of ϵ currently applied ($\epsilon = 34.0$ ev),^[5] we obtain the value 936 ± 35 kev, which agrees well with our measurements. Values of \bar{E}_β which we calculated on the basis of the shape of Y^{90} beta spectra measured on the magnetic spectrograph by Yuasa^[6] and Braden^[7] equal, respectively, 900 and 925 kev with an error of about 5%.

The mean energy of the Y^{90} beta spectrum can also be compared with theoretical values of \bar{E}_β for the unique, first-forbidden beta-transition ($n = 1$, $\Delta J = 2$, the wave function changes sign) with an energy of 2275 kev. Our experimental value is closest to the value of \bar{E}_β computed using the form factors $S = (W_0 - W)^2 L_0 + 9L_1$ ^[8] ($\bar{E}_\beta = 928.8$ kev) and $S = (W^2 - 1) + \lambda(p) \times (W_0 - W)^2$ ^[9] ($\bar{E}_\beta = 929.2$ kev). The form factor $S = (W^2 - 1) + (W_0 - W)^2$,^[10] which does not take into account the influence of the Coulomb field of the nucleus on the emitted electron, yields a somewhat higher value: $\bar{E}_\beta = 942.6$ kev.

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¹N. S. Shimanskaya, JETP **31**, 393 (1956), Soviet Phys. JETP **4**, 355 (1957).

²G. Herrmann and F. Strassmann, Z. Naturforsch. **11a**, 946 (1956).

³N. S. Shimanskaya, Trudy, Radium Inst. **9**, 126 (1959).

⁴R. Caswell, Phys. Rev. **86**, 82 (1952).

⁵J. Boag, Paper at the Symposium on the Methodology of Measurements of Ionizing Radiations, Rome, 1958.

⁶Yuasa, Laberrigue-Frolow, and Feuvrais, Comp. rend. **242**, 2129 (1956).

⁷Braden, Slack, and Shull, Phys. Rev. **75**, 1964 (1949).

⁸K. Siegbahn, Beta- and Gamma-Ray Spectroscopy, Amsterdam, 1955, Russ. Transl. Fizmatgiz 1959, p. 283.

⁹Laslett, Jensen, and Paskin, Phys. Rev. **79**, 412 (1950).

¹⁰E. J. Konopinski and G. E. Uhlenbeck, Phys. Rev. **60**, 308 (1941).

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