

### Isentropic Relativistic Gas Flows

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General baroclinic isentropic relativistic gas flows are analyzed. Equations of vorticity, and a nonlinear equation of propagation of sound waves are derived. In the case of barotropic flow, a relativistic generalization of Thompson's theorem is found.

**I**N classical hydrodynamics, one can prove that for a barotropic isentropic gas flow\* the circulation of the velocity around any closed curve moving with the fluid remains constant in time. In such a flow, vorticity of the velocity field can neither be created nor destroyed. If the flow is at one time described by a velocity potential, it retains this property for all time.

We shall prove that analogous theorems\*\* hold in relativistic hydrodynamics. The situation differs only in that the ordinary 3-velocity of classical hydrodynamics must now be replaced not by the relativistic 4-velocity

$$u_i = g_{ik}u^k = g_{ik}dx^k/ds$$

( $s$  being the proper time), but by the "pseudo-velocity"

$$v_i = Ju_i, \tag{1}$$

where  $J = (w/\rho)$  is the relativistic heat content per unit of rest-energy. Here  $w$  is the relativistic heat-content per unit of proper volume, and  $\rho$  is the rest-energy per unit proper volume, i.e., the rest-energy which the gas would have at absolute zero temperature. The dimensionless state-parameter  $J$  is always greater than unity.

By  $S$  we denote entropy per unit of rest-energy. In relativistic thermodynamics<sup>1</sup>,  $S$  satisfies the equation

$$TdS = dJ - dp/\rho \tag{2}$$

\* By isentropic we mean a flow in which the entropy of each small element of the gas remains constant in time; by barotropic we mean a flow in which the entropy per unit mass is the same for all elements.

\*\* Khalatnikov<sup>2</sup> was the first to investigate relativistic potential flows. We shall here discuss in greater detail flows possessing vorticity.

where  $p$  is the pressure. We make no assumptions about the properties of the gas. We assume an equation of state  $p = p(\rho, T)$ , and a dependence of the internal energy-density  $e$  on  $\rho$  and  $T$ , these relations being completely arbitrary, subject only to the laws of relativistic thermodynamics and to the identity  $w = e + p$ . We further introduce the dimensionless relativistic sound-velocity

$$\bar{a} = a/c = \sqrt{(\partial p/\partial e)_S} \tag{3}$$

where  $c$  is the velocity of light.

We shall prove that, when the ordinary velocity is replaced by the pseudovelocity  $v_i$ , the relativistic theory gives a system of equations for  $v_i$  and  $S$  completely analogous to the equations of classical hydrodynamics. We carry through the analysis for the case of rectilinear coordinates in special relativity, i.e., assuming  $g_{ik}$  constant. The transition to general relativity can be made in the usual way, by rewriting the equations in a form which is invariant under general coordinate transformations.

The energy-momentum equations are\*

$$\partial T_i^k / \partial x_k = 0, \quad T_i^k = wu_iu^k - \delta_i^k p \tag{4}$$

and the equation of conservation of mass is

$$\partial(\rho u^k) / \partial x_k = 0. \tag{5}$$

Using Eq. (5) we reduce Eq. (4) to the form

$$\rho u^k \partial v_i / \partial x_k = \partial p / \partial x_i \tag{6}$$

(the relativistic Euler equations), and hence, by means of Eq. (2) to the form

$$u^h \partial v_i / \partial x_h = \partial J / \partial x_i - T \partial S / \partial x_i. \tag{7}$$

\* The sign of the tensor  $g_{ik}$  is chosen so that the differential of proper times is  $ds = (g_{ik} dx^i dx^k)^{1/2}$ .

Because of the identity

$$g^{hi}u_h u_i = 1$$

the expression (1) for the pseudovelocity implies

$$g^{hi}v_h v_i = J^2. \tag{8}$$

Differentiation with respect to  $x_i$  gives

$$g^{hi}v_i \partial v_h / \partial x_i = J \partial J / \partial x_i$$

or

$$u^h \partial v_h / \partial x_i = \partial J / \partial x_i. \tag{9}$$

Subtracting Eq. (9) from Eq. (7), we find

$$u^k \left( \frac{\partial v_i}{\partial x_k} - \frac{\partial v_k}{\partial x_i} \right) = -T \frac{\partial S}{\partial x_i}. \tag{10}$$

These are the well-known vorticity equations of classical hydrodynamics, with the curl of the velocity replaced by the curl of the pseudovelocity. Multiplying Eq. (10) by  $u^i$  and contracting, we obtain immediately the equation of conservation of particle entropy

$$dS / ds = u^i \partial S / \partial x_i = 0. \tag{11}$$

We proceed to transform the equation of continuity (5). Putting  $u^k = v^k / J$ , we find

$$\frac{\rho}{J} \frac{\partial v^k}{\partial x_k} + J \frac{d(\rho / J)}{ds} = 0.$$

From Eq. (11) and (2) we deduce

$$\frac{d(\rho / J)}{ds} = \frac{\rho}{J^2} (\bar{a}^{-2} - 1) \frac{dJ}{ds}, \tag{12}$$

and so the continuity equation takes the form

$$(dv^k / \partial x_k) + (\bar{a}^{-2} - 1) dJ / ds = 0. \tag{13}$$

From Eq. (8) we have

$$J dJ / ds = v^i dv_i / ds$$

or

$$dJ / ds = u^i dv_i / ds = u^i u^k dv_i / dx_k. \tag{14}$$

Therefore, Eq. (13) becomes

$$[g^{ik} + (\bar{a}^{-2} - 1) u^i u^k] dv_i / \partial x_k = 0. \tag{15}$$

This equation, when both flow and sound velocities

are small, reduces to the classical equation of propagation of sound, including the effect of wind-velocity. In terms of general coordinates, and in general relativity, the corresponding equation is

$$[g^{ik} + (\bar{a}^{-2} - 1) u^i u^k] \times [(\partial v_i / \partial x_k) - \Gamma_{ik}^l v_l] = 0. \tag{16}$$

The whole system of hydrodynamical equations is contained in Eqs. (10) and (15).

The theorem governing the circulation around a line moving with the fluid in a barotropic flow is obtained as follows. In a barotropic flow Eq. (7) gives

$$u^k \partial v_i / \partial x_k = \partial J / \partial x_i \tag{17}$$

or  $\partial v_i / ds = \partial J / \partial x_i,$

where  $d$  means differentiation along the path of a fluid-element. Let  $\delta$  denote differentiation along the line around which we are considering the circulation of pseudovelocity

$$\Gamma = \oint v_i \delta x^i. \tag{18}$$

Then

$$\begin{aligned} d\Gamma / ds &= \oint [(dv_i / ds) \delta x^i + v_i \delta dx^i / ds] \\ &= \oint [(\partial J / \partial x_i) \delta x^i + v_i \delta u^i] \\ &= \oint [\delta J + 1/2 J \delta (u_i u^i)] = \oint \delta J = 0, \end{aligned}$$

showing that  $\Gamma$  remains constant in time. Strictly speaking, we must say that the circulation of pseudovelocity around a fluid line is equal in two successive positions, if each fluid element along the line has lived through the same interval of proper-time in moving from the earlier to the later position. From this theorem follows the impossibility of creating or destroying vortices of pseudovelocity in a barotropic flow.

If in a barotropic flow the pseudovelocity is derived from a potential ( $v_i = d\varphi / dx_i$ ), then the relativistic Euler equations follow automatically. In this case

$$g^{ik} \frac{\partial \varphi}{\partial x_i} \frac{\partial \varphi}{\partial x_k} = J^2,$$

and so

$$g^{ik} \frac{\partial \varphi}{\partial x_i} \frac{\partial^2 \varphi}{\partial x_k \partial x_l} = J \frac{\partial J}{\partial x_l} = \frac{J}{\rho} \frac{\partial p}{\partial x_l}$$

which gives

$$u^k \partial v_k / \partial x_l = \rho^{-1} \partial \rho / \partial x_l,$$

i.e., Eq. (6). In a potential flow, the equation of sound-propagation (15) becomes

$$[g^{ik} + (\bar{a}^{-2} - 1) u^i u^k] \partial^2 \varphi / \partial x_i \partial x_k = 0. \quad (19)$$

In a barotropic flow,  $\bar{a}$  depends only on  $J$ , and hence, by Eq. (8) on the pseudovelocity; therefore,

Eq. (19) contains only the potential  $\varphi$  and its derivatives, and the entire problem in this case reduces to the solution of the single equation (19).

<sup>1</sup> L. D. Landau and E. M. Lifshitz, *Mechanics of Continuous Media*, 2nd Edition, Moscow, 1954.

<sup>2</sup> I. M. Khalatnikov, J. Exptl. Theoret. Phys. (U.S.S.R.) 27, 529 (1954).

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